# Planetary Crater Detection and Registration Using Marked Point Processes, Multiple Birth and Death Algorithms, and Region-based Analysis 

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## Introduction

## Crater Detection

- Marked Point Process Model
- Energy Function
- Multiple Birth and Death Algorithm
- Region-of-Interest Approach
- Experimental Results


## Image Registration

- 2-step Approach
- Experimental Results

Conclusion


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## Int roduction

## Need for automated methods for image registration

$\left.\begin{array}{|c|}\begin{array}{|c}\text { Launch of several } \\ \text { planetary missions }\end{array} \\ \hline \begin{array}{c}\text { Design of new and } \\ \text { powerful sensors }\end{array} \\ \hline\end{array}\right\}$


## Objective

- Crater detection in planetary images
- Development of an image registration method based on the extracted features



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## Marked Point Processes

## Crater detection based on a marked point process (MPP) model

MPP: Stochastic Process $\xrightarrow{\text { Realizations }} \begin{gathered}\text { Configurations of objects, each } \\ \text { described by a marked point }\end{gathered}$

## Mathematical Formulation

A point process $X$, defined over a bounded subset $P$ of $\mathbb{R}^{2}$ maps from a probability space to a configuration of points in $P$.

Realizations of the process $X$ are random configurations $x$ of points, $x=\left\{\boldsymbol{x}_{1}, \ldots, x_{n}\right\}$, where $\boldsymbol{x}_{i}$ is the location of the $i^{\text {th }}$ point in the image plane ( $\boldsymbol{x}_{i} \in P$ )

A configuration of an MPP consists of a point process whose points are enriched with additional parameters, called marks and aimed at parameterizing objects linked to the points.

Bayesian approach: Maximum a posteriori (MAP) rule to fit the model to the image is equivalent to minimizing an energy function (computationally challenging)

## Marked Point Process for Crater Detection



## Crater Detection - Energy Function

Energy function of the configuration $X=\left\{\boldsymbol{x}_{i}, \boldsymbol{x}_{2}, \ldots, \boldsymbol{x}_{n}\right\}$ wrt the extracted set $C$ of contour pixels (Canny):

$$
U(X \mid C)=U_{P}(X)+U_{L}(C \mid X)
$$

## Prior

Repulsion coefficient based on the overlapping of the ellipses (overlapping craters are quite unlikely)

$$
U_{P}(X)=\frac{1}{n} \sum_{x_{i} \wedge x_{j}>0} \frac{\boldsymbol{x}_{i} \wedge \boldsymbol{x}_{j}}{\boldsymbol{x}_{i} \vee \boldsymbol{x}_{j}}
$$

$x_{i} \vee x_{j}=$ area of union of ellipses $x_{i}$ and $x_{j}$ $x_{i} \wedge x_{j}=$ area of intersection of $x_{i}$ and $x_{j}$

## Likelihood

Two terms, one based on a correlation measure, the other based on a distance measure (fit between contours and realization of $X$ )

$$
U_{L}(C \mid X)=\sum_{i=1}^{n}\left[\frac{d_{\mathcal{H}}\left(x_{i}^{0}, C\right)}{n a_{i}}-\frac{\left|x_{i}^{0} \cap C\right|}{|C|}\right]
$$

$x_{i}^{0}=$ set of pixels corresponding to ellipse $x_{i}$ in the image plane $d_{\mathcal{H}}\left(x_{i}^{0}, C\right)=$ Hausdorff distance between ellipse $\boldsymbol{x}_{i}$ and the contours:


$$
d_{\mathcal{H}}(A, B)=\max \left\{\sup _{\alpha \in A} \inf _{\beta \in B} d(\alpha, \beta) ; \sup _{\beta \in B} \inf _{\alpha \in A} d(\alpha, \beta)\right\}
$$

Classical distance between sets $d(A, B)=0$

## Crater Detection - Energy Minimization

Markov chain Monte Carlo-type method Simulated Annealing scheme

Markov chain sampled by a multiple birth and death
(MBD) algorithm


## MBD - Birth and Death Steps

## Birth Step

For each pixel $s$ in the image, compute the birth probability as $\min \{\delta \cdot B(s), 1\}$, where:

$$
B(s)=\frac{b(s)}{\sum_{s} b(s)}
$$

$b(s)$ is the birth map computed from the contour map using generalized Hough transform and Gaussian filtering


## Death Step

For each ellipse $x_{i}$ in the configuration, compute the death probability as $d\left(\boldsymbol{x}_{i}\right)$ :

$$
d\left(\boldsymbol{x}_{i}\right)=\frac{\delta \cdot a\left(\boldsymbol{x}_{i}\right)}{1+\delta \cdot a\left(\boldsymbol{x}_{i}\right)}
$$

$a\left(\boldsymbol{x}_{i}\right)=\exp \left[-\beta\left(U_{L}\left(X \backslash\left\{\boldsymbol{x}_{i}\right\} \mid C\right)-U_{L}(X \mid C)\right)\right]$

## Crater Detection - Region Based Approach

| Region-Based Approach | Why? | - MBD is computationally heavy |
| :---: | :---: | :---: |
|  |  | Computational burden increases with image size |

## Region Based Flowchart and Example



## Crater Detection - Data Sets

- 6 THEMIS (Thermal Emission Imaging System) images, TIR, 100 m resolution, Mars Odissey mission
- 7 HRSC (High Resolution Stereo Color) images, VIS, $\sim 20 \mathrm{~m}$ resolution, Mars Express mission
- Image sizes from $1581 \times 1827$ to $2950 \times 5742$ pixels

| Data | $D=\frac{T P}{T P+F N}$ | $B=\frac{F P}{T P}$ | $Q=\frac{T P}{T P+F P+F N}$ |
| :---: | :---: | :---: | :---: |
| Avg on all THEMIS | 0.91 | 0.10 | 0.83 |
| Avg on all HRSC | 0.89 | 0.06 | 0.85 |
| Avg on all images | 0.90 | 0.09 | 0.84 |

## Crater Detection - Results



Crater geometric properties extracted by the proposed method

| Crater | $\boldsymbol{C}=\left(\boldsymbol{x}_{\mathbf{0}}, \boldsymbol{y}_{\mathbf{0}}\right)$ | Semi-axes $(\boldsymbol{a}, \boldsymbol{b})$ | Orientation $\boldsymbol{\theta}$ |
| :---: | :---: | :---: | :---: |
| Crater 1 | $(139,393)$ | $(35,33)$ | $64^{\circ}$ |
| Crater 2 | $(258,756)$ | $(51,50)$ | $115^{\circ}$ |
| Crater 3 | $(343,23)$ | $(13,12)$ | $180^{\circ}$ |
| Crater 4 | $(591,215)$ | $(19,18)$ | $31^{\circ}$ |
| Crater 5 | $(919,157)$ | $(15,14)$ | $106^{\circ}$ |

## HRSC Sensor

THEMIS Sensor


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## Image Registration-2-Step Optimization



Why a 2-step Optimization?

## Feature-based registration

- Min Hausdorff distance ( $d_{\mathcal{H}}$ ) between extracted craters through genetic algorithm
- Fast but sensitive to accuracy of crater maps


## Area-based registration

- Max Mutual Information (MI) through genetic algorithm
- Highly accurate but computationally heavy


Fast registration based on extracted craters $\rightarrow \widetilde{\boldsymbol{p}}$


Refinement: registration based on mutual information in a neighborhood of $\widetilde{\boldsymbol{p}} \rightarrow \boldsymbol{p}^{*}$

Transformation found for an interactively selected region of interest $\rightarrow p_{B}^{*}$

$$
p_{B}=\left(t_{x}, t_{y}, \theta, k\right)
$$



Transformation derived for the entire Image $\rightarrow p_{A}^{*}$

$$
p_{A}=\left(T_{x}, T_{y}, \beta, \alpha\right)
$$



Superposition of Reference and Input


## Image Registration - Data Sets

## Semi-simulated image pairs

20 pairs composed of one real THEMIS or HRSC image and of an image obtained by applying a synthetic transform and AWGN

Quantitative validation with respect to the true transform (RMSE)


## Real multi-temporal image pairs

Real multi-temporal pair of LROC (Lunar Reconnaissance Orbiter Camera) images

100m resolution
Only qualitative visual analysis is available, as no ground truth is available


## Registration Results with Semi-synthet ic

 DataRGB of original non-
registered data


| Data set | RMSE [pixel] |  | $p_{G T}$ | $\left(7.05,35.91,0.18^{\circ}, 1.071\right)$ | $\left(76.59,19.96,2.17^{\circ}, 1.031\right)$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| THEMIS (10 data sets) | 0.31 |  | $p^{*}$ | $\left(7.04,35.92,0.19^{\circ}, 1.071\right)$ | $\left(76.41,20.06,2.18^{\circ}, 1.031\right)$ |
| HRSC (10 data sets) | 0.22 |  | RMSE $1^{\text {st }}$ Step | 0.79 | 0.51 |
| Average (20 data sets) | 0.26 |  | RMSE 2 $2^{\text {nd }}$ Step | 0.16 | 0.33 |

## Registration Results with Real Data

0

> Visually accurate matching between reference and registered images in the real multitemporal data set

Checkerboard representation of the registered images (zoom on details)



Visually accurate matching between reference and registered images in the real multitemporal data set


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## Conclusions and Future Developments

## Conclusions

- Accurate crater maps, useful for both image registration and planetary science, were obtained from data from different sensors.
- Higher accuracy as compared to previous work on crater detection (not shown for brevity)
- Reduced time for convergence thanks to a region-based approach
- Sub-pixel accuracy and visual precision in registration: effectiveness of the proposed 2 -step registration method


## Future Developments

- Test in conjunction with a parallel implementation (e.g. computer cluster)
- Validation with multi-sensor real images
- Extension to other applications requiring the extraction of ellipsoidal or circular features, e.g. optical Earth observation images or medical images


## Short Bibliography

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## Appendix

## MBD - Birth Step

For each pixel in the image compute the Birth Probability as $\min \{\delta \cdot B(s), 1\}$, where:

$$
B(s)=\frac{b(s)}{\sum_{s} b(s)}
$$

Being $b(s)$ the Birth Map computed from the Canny Contour Map


## MBD - Death Step

For each ellipse $x_{i}$ in the configuration compute the Death Probability as $d\left(x_{i}\right)$, where

$$
d\left(x_{i}\right)=\frac{\delta \cdot a\left(x_{i}\right)}{1+\delta \cdot a\left(x_{i}\right)} \quad \text { and } \quad a\left(x_{i}\right)=e^{-\beta\left(U_{L}\left(\left\{x \backslash x_{i}\right\} \mid I_{g}\right)-U_{L}\left(x \mid I_{g}\right)\right)}=e^{\beta \cdot U_{L}^{i}\left(x_{i} \mid I_{g}\right)}
$$

The complete Flowchart of the Death Step is as follows:


## Similarit y Measures

## Hausdorff Distance

$$
\text { Similarity }=\operatorname{mean}_{c}\left\{\sum_{i=1}^{N^{c}} \sum_{t=1}^{P}\left[d_{H}\left(\underline{x}_{i}^{c}, \underline{x}_{t}\right)\right]\right\}
$$

$\mathrm{c}=$ craters in Input Image
$\mathrm{N}^{\mathrm{c}}=\operatorname{sum}$ (pixels in crater c in Input Image)
$\mathrm{P}=\operatorname{sum}$ (craters'border pixels in Ref Image)
$\underline{x}_{i}^{c}=$ coord of pixel i in crater $c$ in Input Image
$\underline{x}_{t}=$ coord of pixel $t$ in Ref Image's craters


## Mutual Information

$M I(X, Y)=\sum_{x \in X} \sum_{y \in Y} p_{X, Y}(x, y) \log \left(\frac{p_{X, Y}(x, y)}{p_{X}(x) p_{Y}(y)}\right)$
$X$ : pixel intensity in Reference Image
$Y$ : pixel intensity in Input Image
$p_{X}(x)$ : probability density function (pdf) of $X$ $p_{Y}(y)$ : probability density function (pdf) of $Y$ $p_{X, Y}(x, y)$ : joint pdf of $X$ and $Y$



## RST Transformation

## Rotation - Scale - Translation Transformation

Transformation vector

$$
p=\left(t_{x}, t_{y}, \theta, k\right)
$$

$\left\{t_{x}, t_{y}\right\}$ : Translations in $x$ and $y$
$\theta$ : Rotation angle
Matrix Formulation

$$
T_{p}(x, y)=\left(\begin{array}{ccc}
k \cos (\theta) & k \sin (\theta) & t_{x} \\
-k \sin (\theta) & k \cos (\theta) & t_{y}
\end{array}\right)\left[\begin{array}{l}
x \\
y \\
1
\end{array}\right]
$$

$k$ : Scaling Factor

Original


Rotation


Scaling


Translation


## Region of Interest Approach



Expressing the transformation in Matrix Form

$$
\begin{array}{ll}
\text { From the image } & T_{p_{A}}=\left(\begin{array}{ccc}
\alpha \cos (\beta) & \alpha \sin (\beta) & T_{x} \\
-\alpha \sin (\beta) & \alpha \cos (\beta) & T_{y}
\end{array}\right): \\
\qquad \begin{array}{lll}
X=x+x_{0} & & \left.T_{p_{A}}, Y\right)=\left(X^{\prime}, Y^{\prime}\right) \\
Y=y+y_{0} & & T_{p_{B}}=\left(\begin{array}{ccc}
k \cos (\theta) & k \sin (\theta) & t_{x} \\
-k \sin (\theta) & k \cos (\theta) & t_{y}
\end{array}\right): \\
T_{p_{B}}(x, y)=\left(x^{\prime}, y^{\prime}\right)
\end{array}
\end{array}
$$

This should also hold

$$
T_{p_{A}}\left(x+x_{0}, y+y_{0}\right)=\left(x^{\prime}+x_{0}, y^{\prime}+y_{0}\right)
$$

Plugging $T_{p_{A}}$ into this equation and replacing

$$
x^{\prime} \text { and } y^{\prime} \text { according to } T_{p_{B}}
$$

Knowing $\alpha=k$ and solving in $P_{1}=(0,0)$ and $P_{2}=\left(-x_{0},-y_{0}\right)$

$$
\left\{\begin{array}{c}
k \cos (\theta) x+k \sin (\theta) y+t_{x}+x_{0}= \\
\alpha \cos (\beta)\left(x+x_{0}\right)+\alpha \sin (\beta)\left(y+y_{0}\right)+T_{x} \\
-k \sin (\theta) x+k \cos (\theta) y+t_{y}+y_{0}= \\
-\alpha \sin (\beta)\left(x+x_{0}\right)+\alpha \cos (\beta)\left(y+y_{0}\right)+T_{y}
\end{array}\right.
$$

$$
p_{A}=\left(\begin{array}{c}
-k \cos (\theta) x_{0}-k \sin (\theta) y_{0}+t_{x}+x_{0} \\
k \sin (\theta) x_{0}-k \cos (\theta) y_{0}+t_{y}+y_{0} \\
\theta \\
k
\end{array}\right)
$$

## Erorr Transformation

Ground Truth Transformation $p_{G T}=\left(t_{x 1}, t_{y 1}, \theta_{1}, k_{1}\right) \rightarrow T_{p_{G T}}(x, y)=Q_{p_{G T}} \cdot[x, y, 1]^{T}$

ComputedTransformation

$$
p=\left(t_{x}, t_{y}, \theta, k\right) \rightarrow T_{p}(x, y)=Q_{p} \cdot[x, y, 1]^{T}
$$

$(x, y) \in$ Image, $\left[x^{\prime}, y^{\prime}, 1\right]^{T}=Q_{P_{e}} \cdot[x, y, 1]^{T}$

$$
\longmapsto\left[\begin{array}{l}
x^{\prime} \\
y^{\prime}
\end{array}\right]=k_{e}\left(\begin{array}{cc}
\cos \left(\theta_{e}\right) & \sin \left(\theta_{e}\right) \\
-\sin \left(\theta_{e}\right) & \cos \left(\theta_{e}\right)
\end{array}\right)\left[\begin{array}{l}
x \\
y
\end{array}\right]+\left[\begin{array}{l}
t_{x e} \\
t_{y e}
\end{array}\right]
$$

$$
\text { RMS Error: } E\left(p_{e}\right)=\sqrt{\frac{1}{A B} \int_{0}^{A} \int_{0}^{B}\left(x^{\prime}-x\right)^{2}+\left(y^{\prime}-y\right)^{2} d x d y}, \quad \alpha=A^{2}+B^{2}
$$

$$
E^{2}\left(p_{e}\right)=\frac{1}{A B} \int_{0}^{A} \int_{0}^{B}\left(k_{e} \cos \left(\theta_{e}\right) x+k_{e} \sin \left(\theta_{e}\right) y+t_{x e}-x\right)^{2}+\left(-k_{e} \sin \left(\theta_{e}\right) x+k_{e} \cos \left(\theta_{e}\right) y+t_{y e}-y\right)^{2} d x d y
$$

$$
E^{2}\left(p_{e}\right)=\frac{\alpha}{3}\left(k_{e}^{2}-2 k_{e} \cos \left(\theta_{e}\right)+1\right)+\left(t_{x e}^{2}+t_{y e}^{2}\right)-\left(A t_{x e}^{2}+B t_{y e}^{2}\right)\left(1-k_{e} \cos \left(\theta_{e}\right)\right)-k_{e}\left(A t_{y e}-B t_{x e}\right) \sin \left(\theta_{e}\right)
$$

